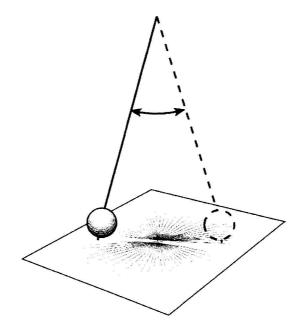


_

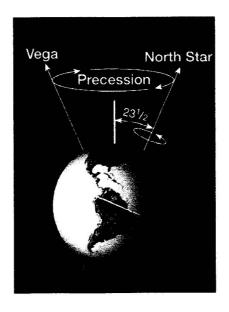
J. MIIIS © 199

Astronomy = 3 (d) (c) S Az Az h h SSE 42 horizon **Common Terms** NW 00 3150 60 157 Degrees II. Gravity Newton's Laws of Gravity Α. 1. All objects possess gravity and will pull all other objects with a certain gravitational force. 2. The mass of an object determines the amount of gravitational force that object possesses. The greater the mass, the greater the gravitational force. 3. The gravitational force between two Objects changes as the distance between change. As distance increases, them arauity decreases.

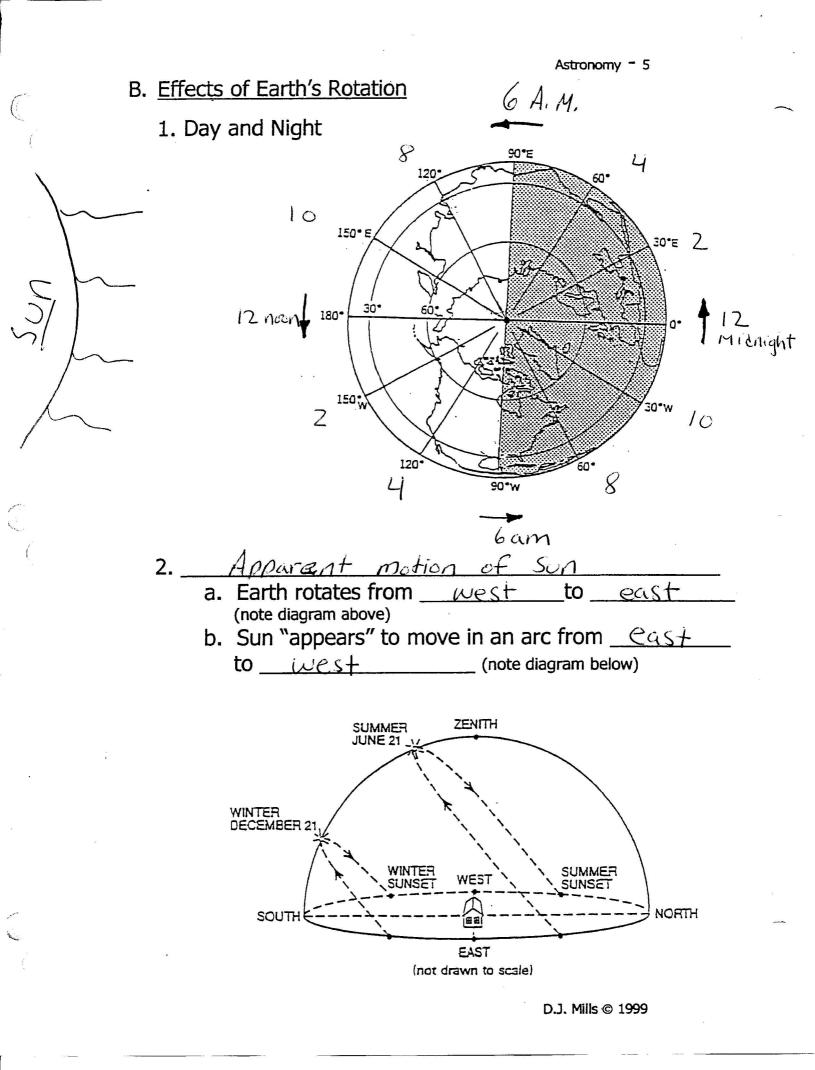
A Foucault Pendulum proves the earth rotates.



Precession - A slow motion of earth's axis.

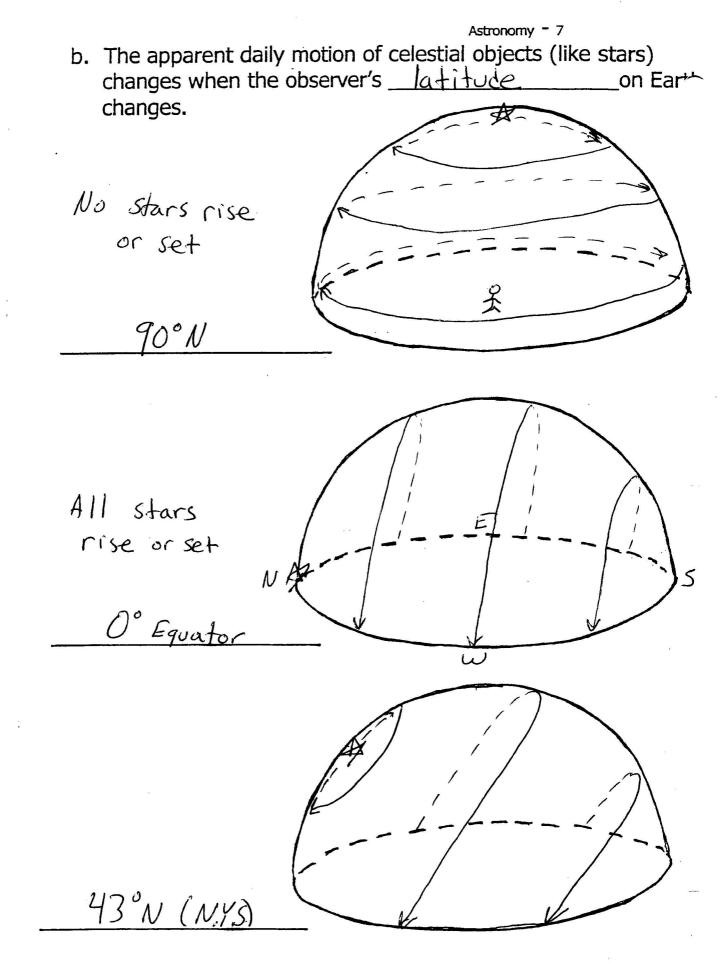


Astronomy = 4 Gravity and Inertia B. 1. Newton's Law of Inertia states that <u>an objects</u> motion will not change unless that object is acted on by an outside force 2. SUN Inertia Gravity lanet Stable orbit a. Inertra - causes a planet to move in a straight line. b. Gravity - pulls a planet towards the sun. Rotation - the spinning of a celestral body. III. (Earth) on an imaginary axis A. Earth's 1. Direction of Rotation: west to east 2. Angular Rate of Rotation: ROTATION THINK - one complete rotation a. <u>360</u>^e degrees b. 24 hours 150 hr RATE = 360° 24655 D.J. Mills @ 1999



Astronomy - 6 3. Apparent daily motion of the stars C Zenith Star X Polaris Stars rise Star Y Observer 1 Stars set N W Horizon (-Ν S N S LOOKING WEST LOOKING EAST Polaris W Ε W Ε LOOKING SOUTH LOOKING NORTH

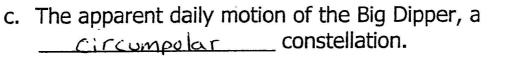
*

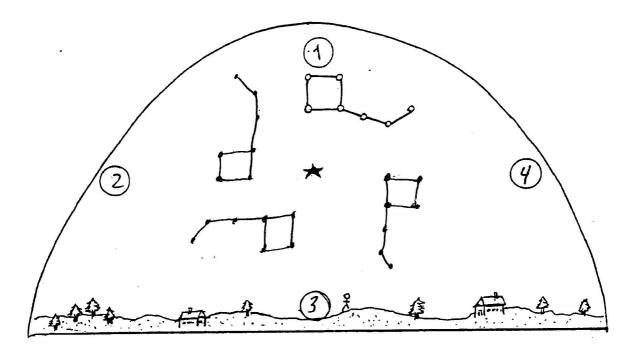


C

(

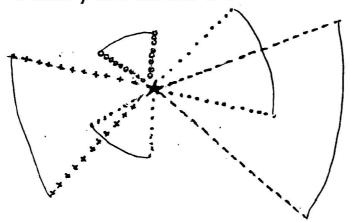
D.J. Mills © 1999

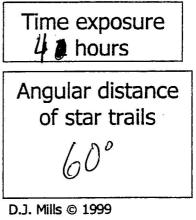


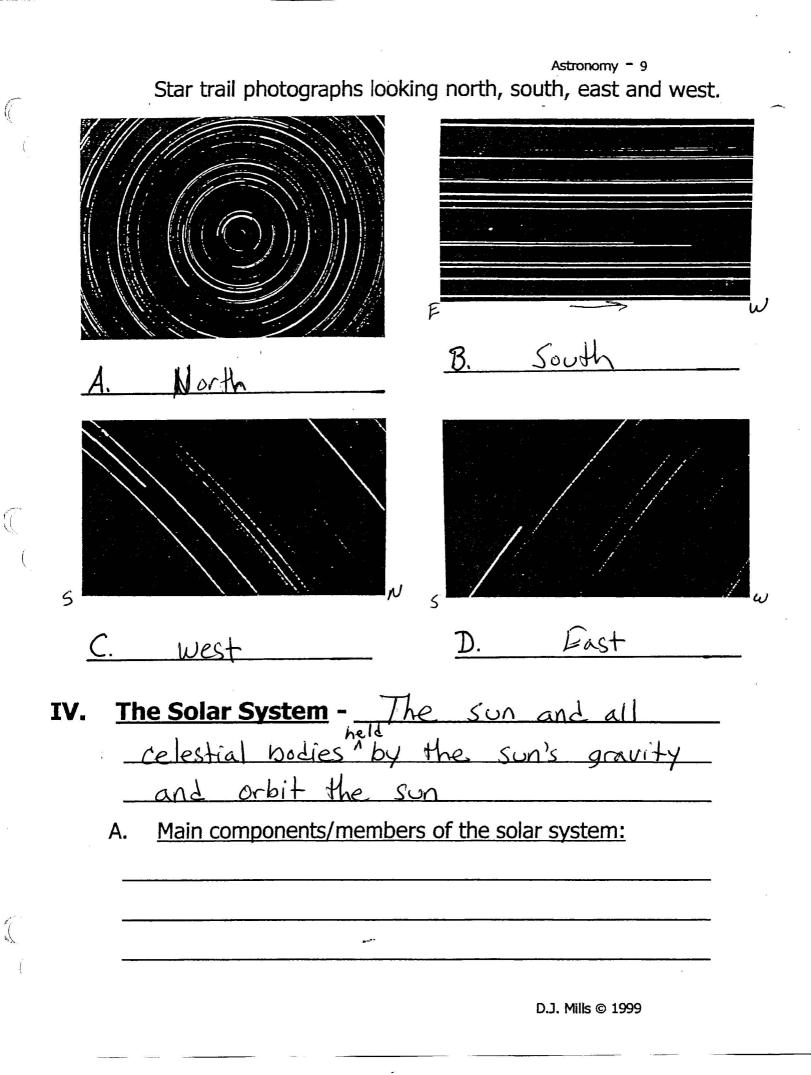


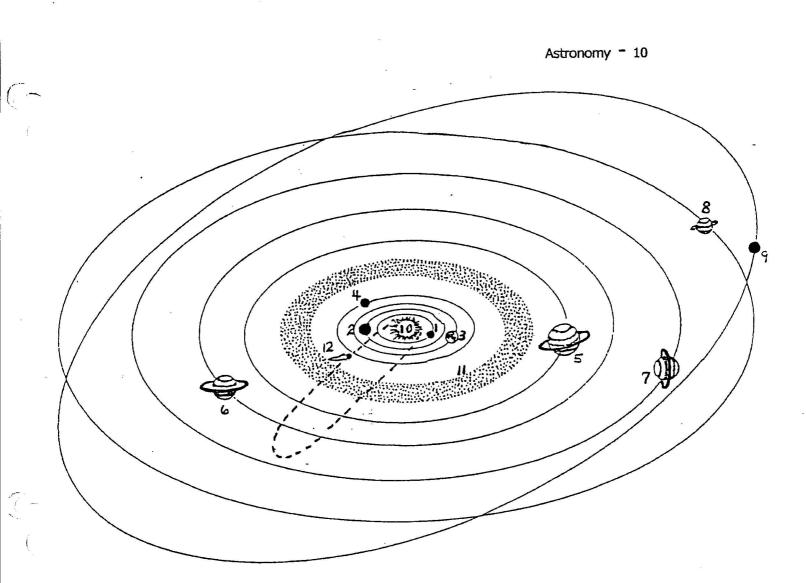
LOCATION OF OBSERVER	N.Y.S.		
DATE	mid Oct		
POSITION OF THE BIG DIPPER	9:00 pm	1	
000	3:00 am	2	
	9:00 am	3	
	3:00 pm	4	

d. 5 + 4 + 1 = 1 = - a time-exposed photographic image that shows the apparent motion of stars; it appears as a blurry line across the film. Time exposure





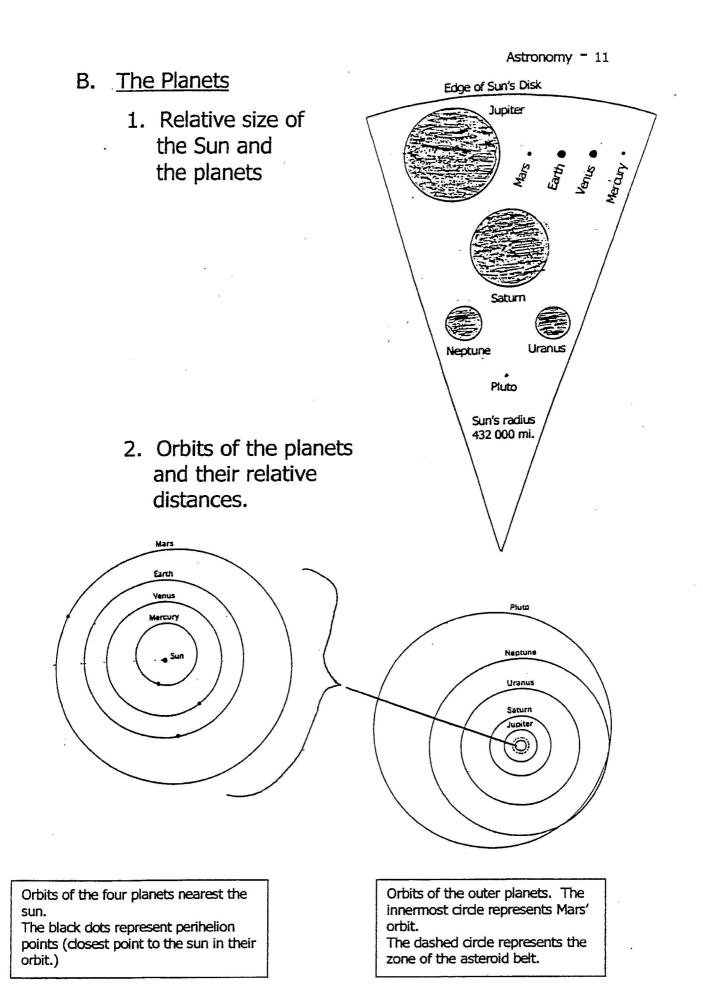




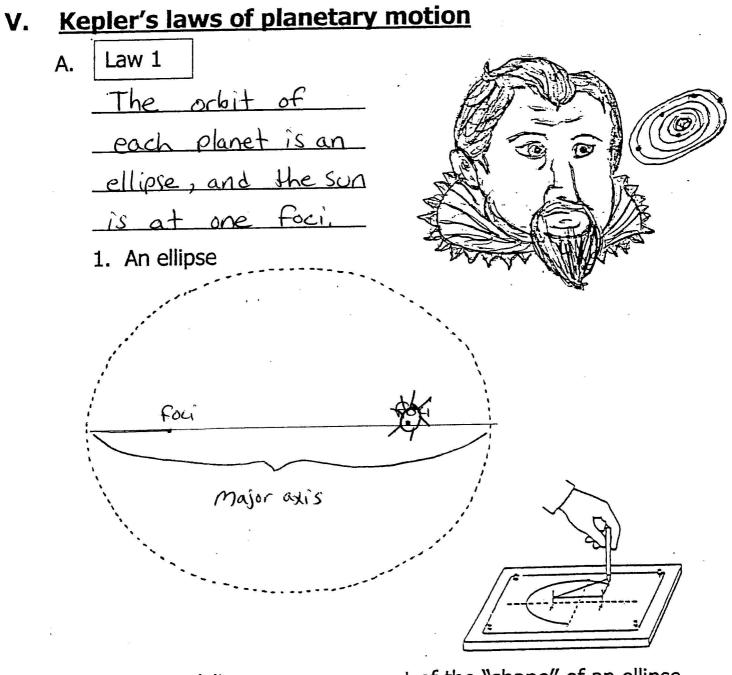
	NAME OF PLANET	SYMBOL	
1	Mercury	<u> </u>) Terrestial
2	Venus	<u>\$</u>	Enner 'Earth
3	Earth	Ø	Like planets"
4	Mars	07	2
5	Jupiter	4	$\left(\right)$
6	Saturn	h	> Cras Giants
7	Uranus	8	Jovian
8	Neptune	Y	
9	Pluto	P]
10 _	Sun		
11 _	Asteroids		
12 _	comet		

•

Ľ,



1



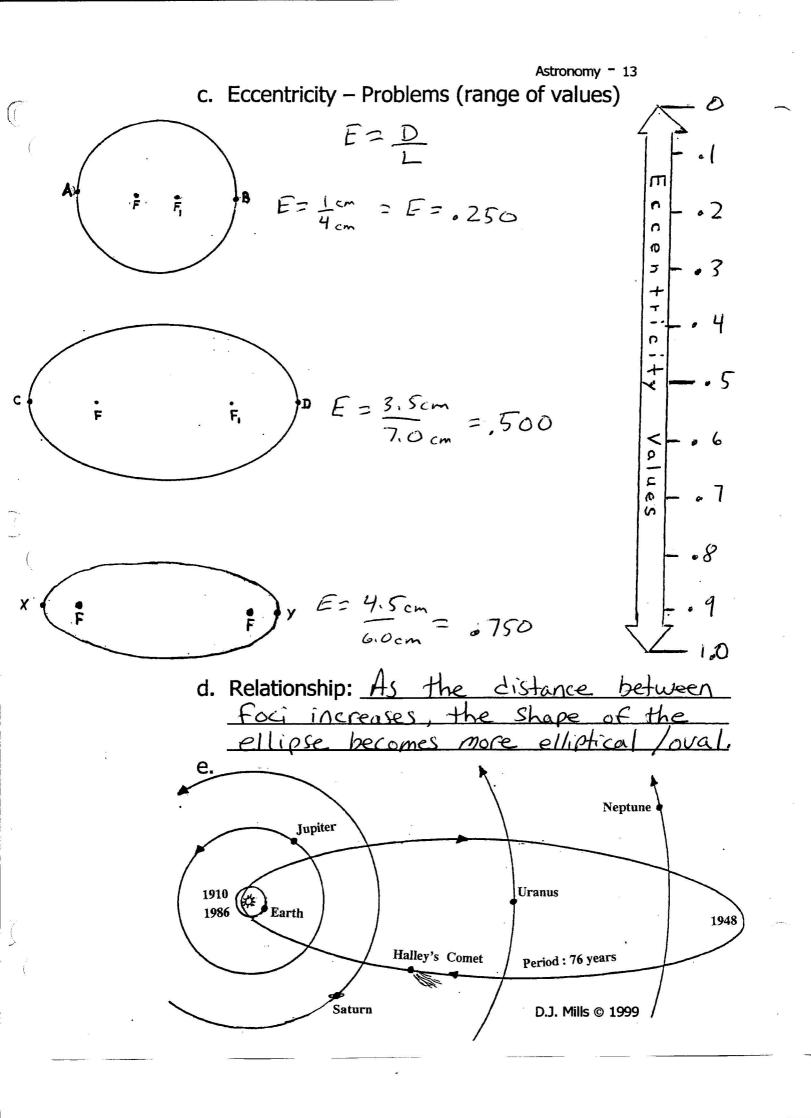
2. Eccentricity – a measurement of the "shape" of an ellipse.

a. Formula:

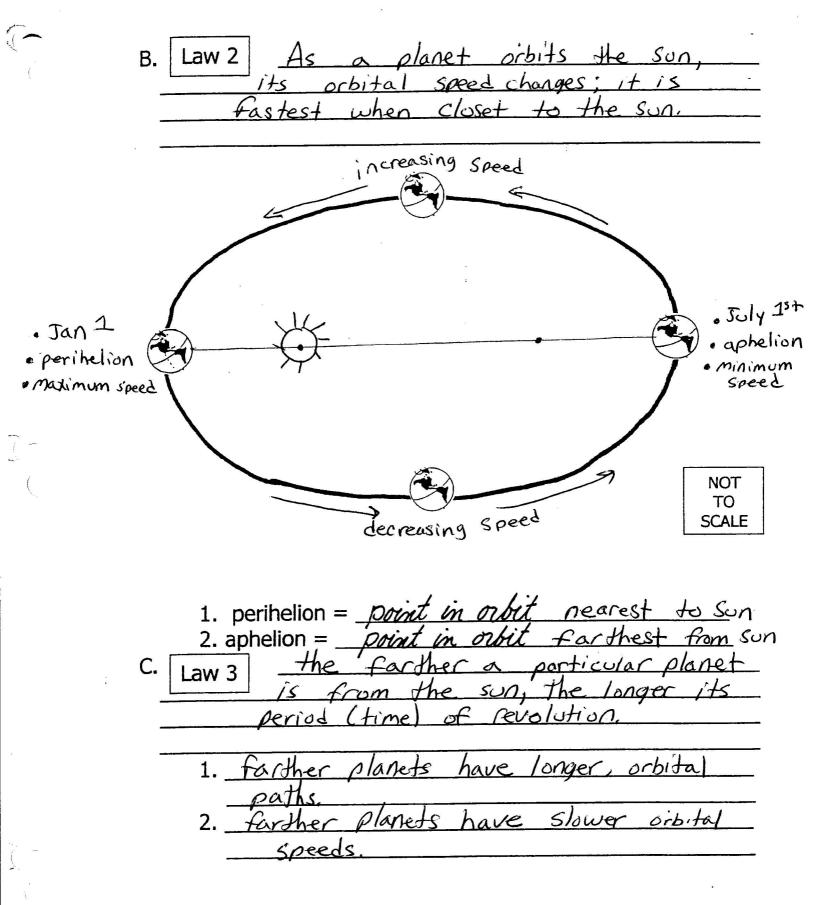
Reference Table page /

b. Sample Problem – based on the ellipse above.

$$E = \frac{6.2 \text{ cm}}{10.5 \text{ cm}} = .5905$$



Astronomy - 14



3. Solar System

4

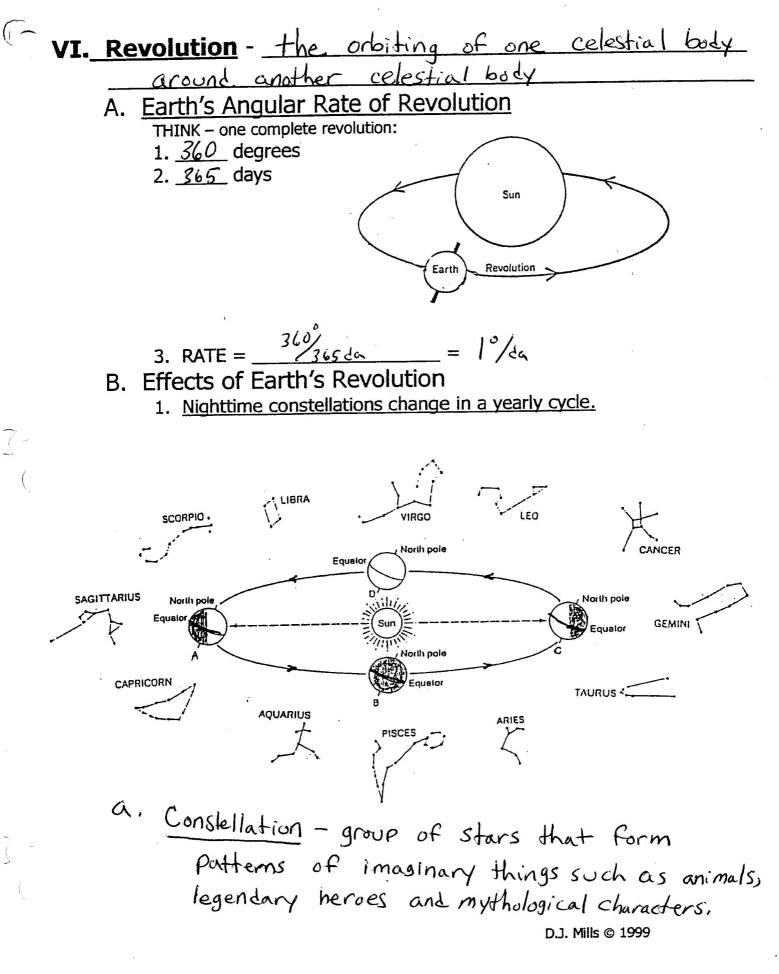
(

) ()

4.				
PLANET	DISTANCE		SPEED	PERIOD OF
from Sun outward	millions of miles	(million km)	mi./sec	REVOLUTION
Mercury	36	57.9	30	88 days
Venus	67	108.2	22	224 days
Earth	93	149.6	19	365 1/4 days
Mars	142	227.9	15	687 days
Jupiter	484	778.4	8	11.86 years
Saturn	887	1,426,7	6.	29.46 years
Uranus	1784	8,871,0	4	84,01 years
Neptune	2795	4,498.3	3,4	164 years
Pluto	3675	149.6	3	247 years

-

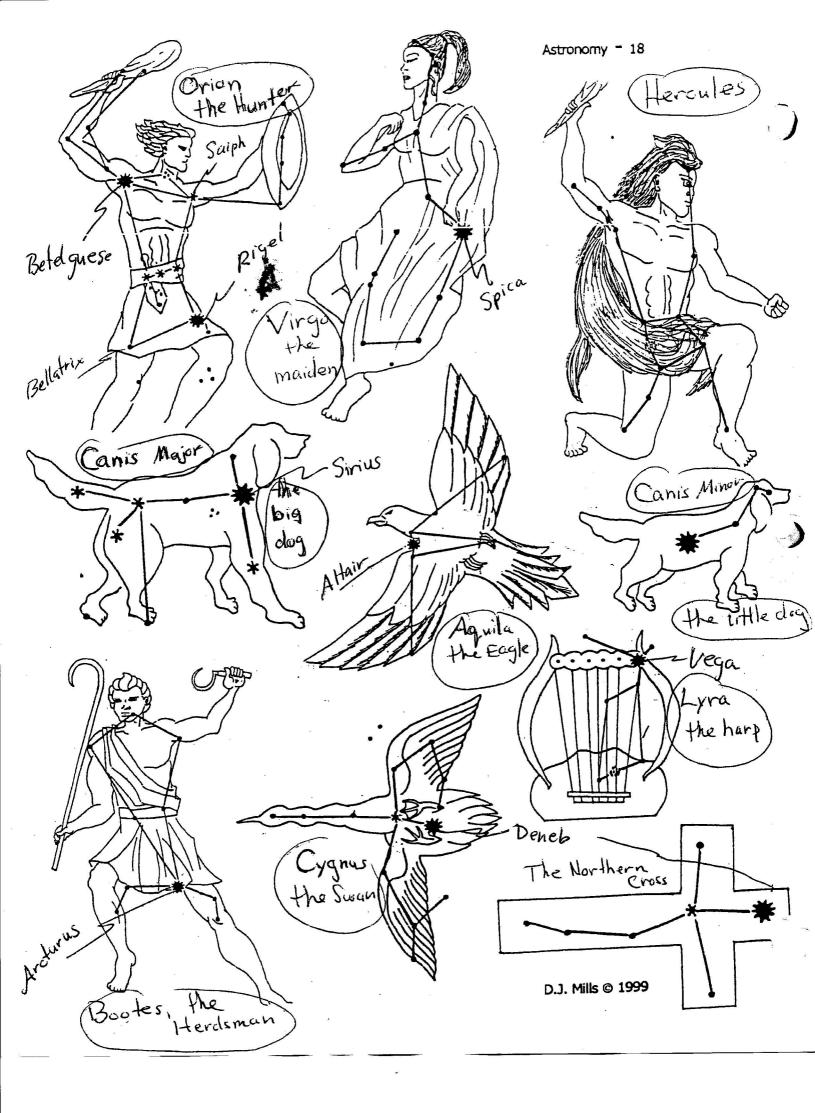
Astronomy - 16



- Astronomy = 17
- b. Zodiac a band of twelve constellations that forms a background for the Sun as seen from the revolving Earth.
- c. Complete the data table below based on the diagram on the preceding page.

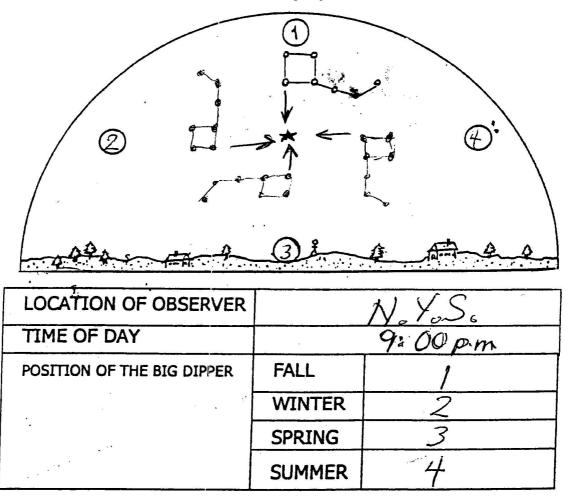
POSITION OF EARTH	SEASON	CONSTELLATIONS VISIBLE AT NIGHT		
Α	Summer	Scorpio, Sagittarius, Capricorn		
В	Fall	Aquarius, Pises, Aries		
C	winter	Taurus, Gemini, Cancer		
D	spring	Leo, Virgo, Libra		
d. <u>The Constellations</u> Antares Antares Gemin Denebula Denebula				
Star Magni $=1^{st}$ $=2^{nd}$ $=3^{rd}/2$	tudes	c, the Lion Corona the northern craw		

.....

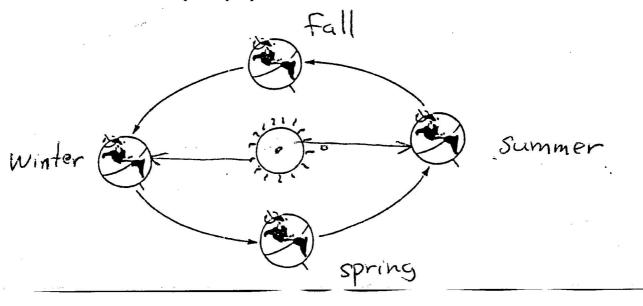


Continued: Effects of Earth's Revolution

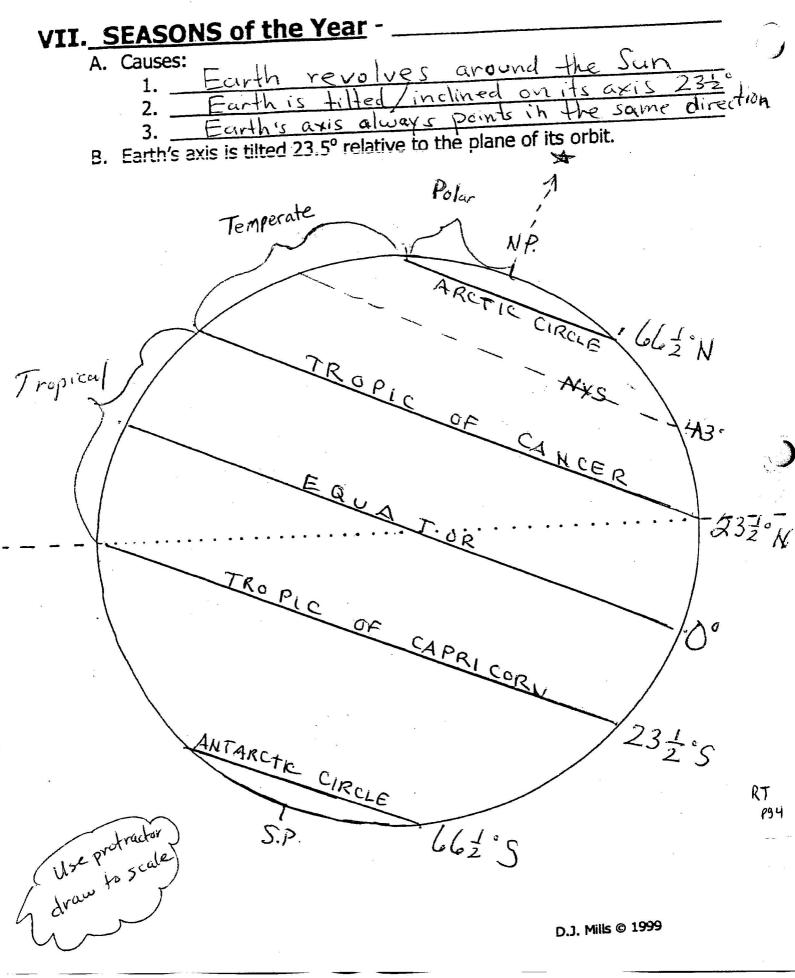
2. Position of the Big Dipper (and other circumpolar constellations) changes position in a yearly cycle.

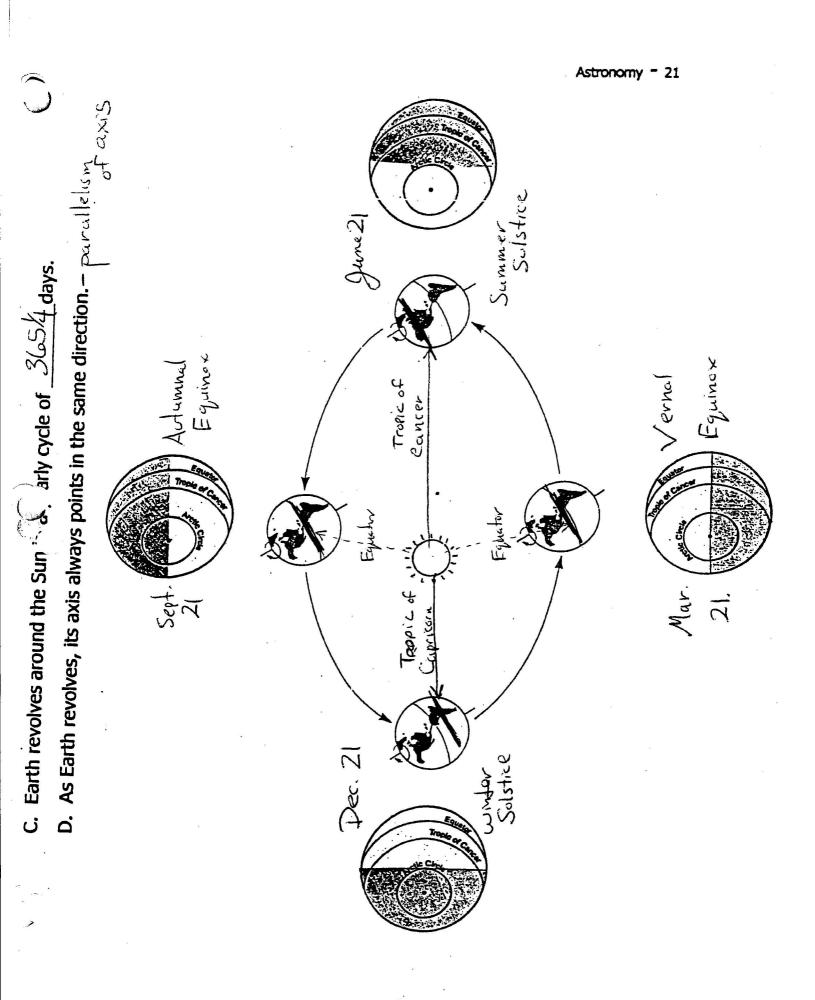


3. Seasons – a yearly cycle

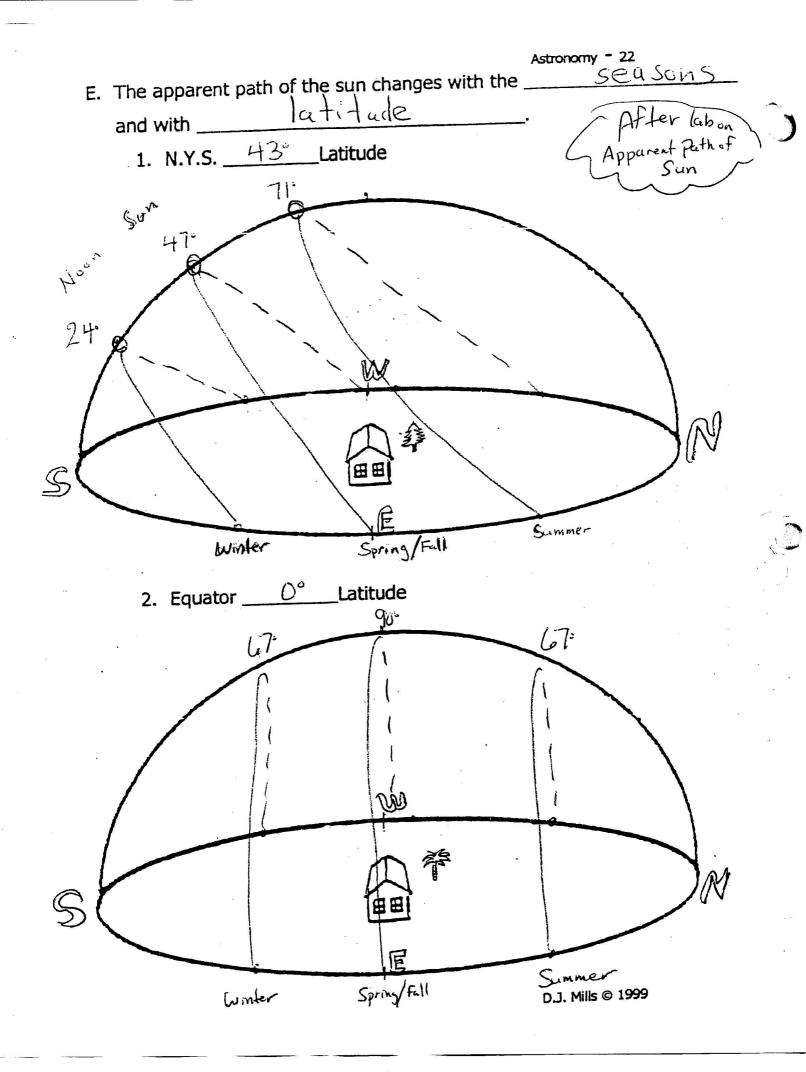


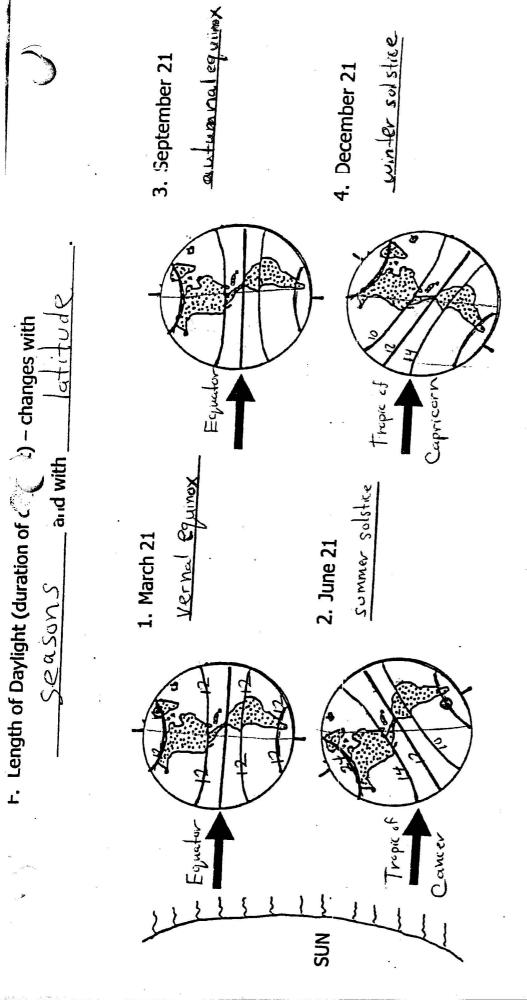
Astronomy - 20



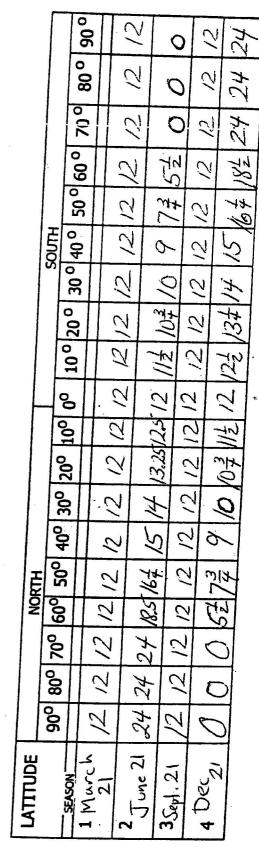


D.1. Milk @ 1999





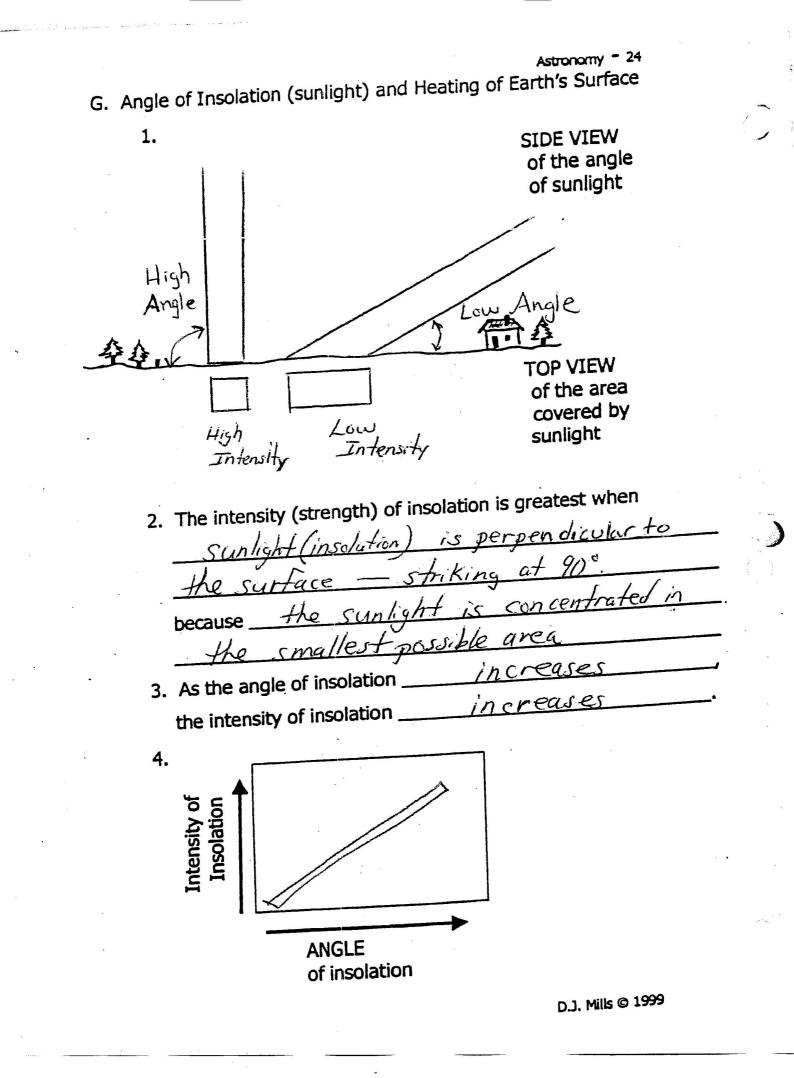
Astronomy = 23

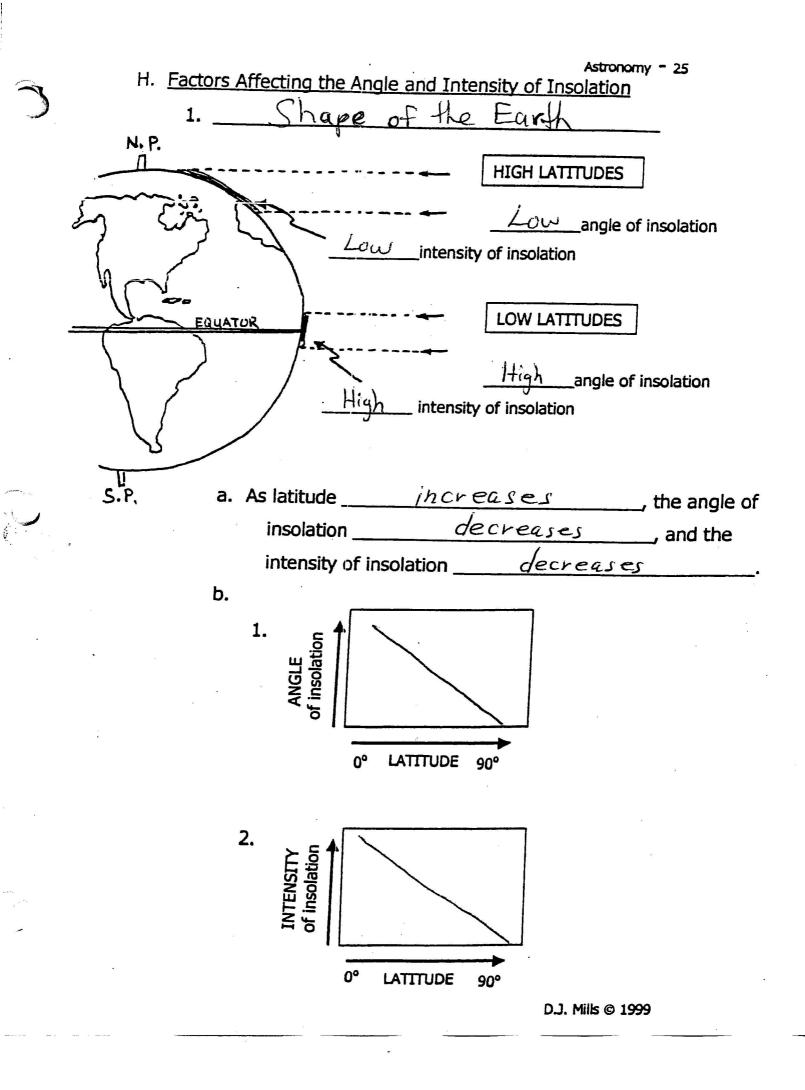


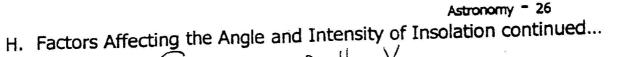
HOURS OF DAYLIGHT

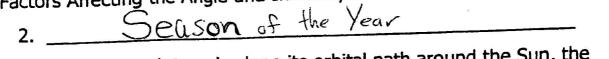
D.J. Mills © 1999

1

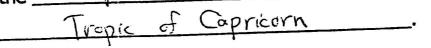




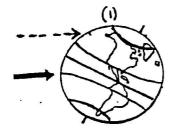


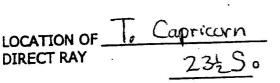


a. As Earth travels along its orbital path around the Sun, the angle of insolation at a given latitude changes with the seasons. This depends on how far a given latitude is from the direct rays of the Sun. The direct rays migrate between and the of Cancer Tropic the



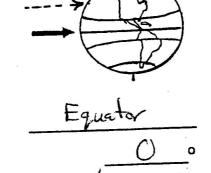
- b. Key: Direct Ray of Sunlight (perpendicular to Earth's surface; 90°)
 - ----→ Ray of Sunlight (striking N.Y.S. at an angle less than 90°) (2)

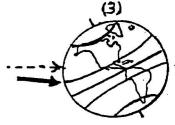


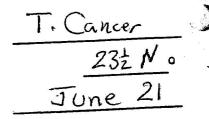


bc.

21







DATE

WINTER SOLSTICE SEASON

> c. Maximum angle of insolation at 12 Noon for mid N.Y.S. (latitude 43°N)

Nar

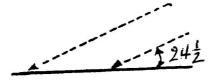
EQUINOXES

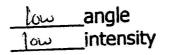
1. Dec. 21

2. Mar/Sept 21



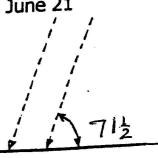
SUMMER SOLSTICE





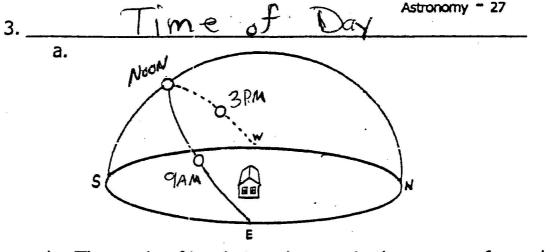


medium angle moderate intensity

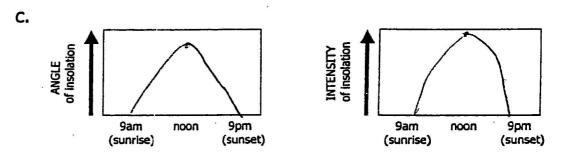


angle high intensity

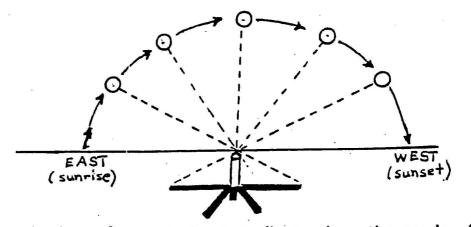
D.J. Mills © 1999



b. The angle of insolation changes in the course of one day. Maximum intensity occurs at N^{OON} .



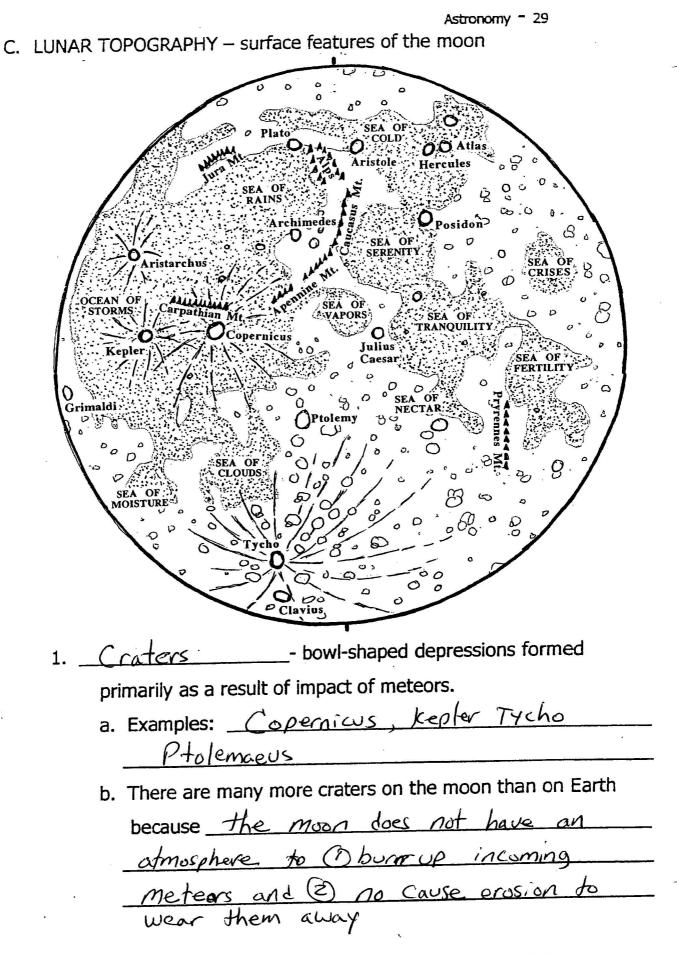
d. Looking south



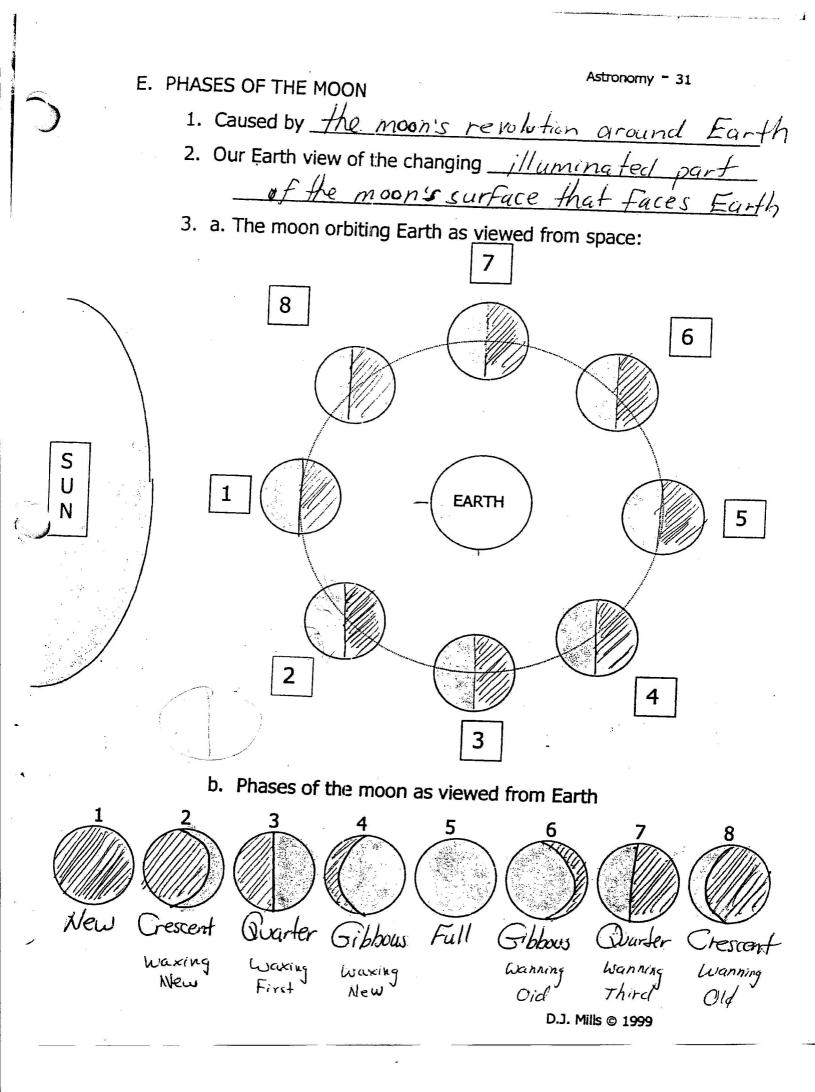
The shadow of a vertical pole indicates how the angle of insolation changes during the day. The higher/greater the angle of insolation, the <u>shorter</u> the shadow, and the <u>greater</u> the intensity of insolation.

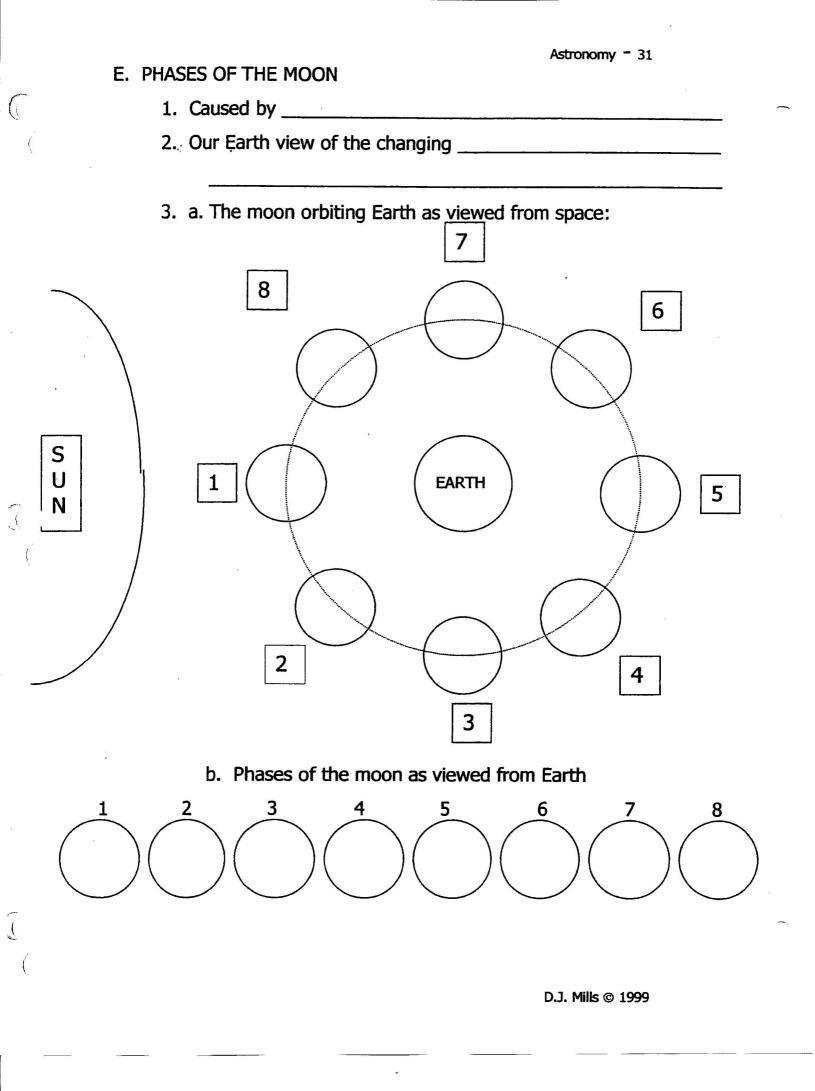
Astronomy - 28

VIII The Moon A. The moon is a Matural satellite of earth 1. Juna - Latin word for the moon 2. ______ - Roman goddess of the moon B. PHYSICAL PROPERTIES OF THE MOON 1. SIZE a. Diameter: 260 miles b. Compared to Earth Diameter of: MOON _ 2160 8000 miles EARTH c. Scale of Size: 2. GRAVITY a. <u>1/6</u> the gravity of Earth b. <u>Smaller - less mass</u> 3. ATMOSPHERE a. ____ virtually none_____ b. Gravity too weak - gases escape out into space 4. TEMPERATURES a. _ 240°F _____ on the lighted side b. $-240^{\circ F}$ on the dark side c. These large temperature extremes or differences exist because the moon dues not have an atmosphere to transfer heat.

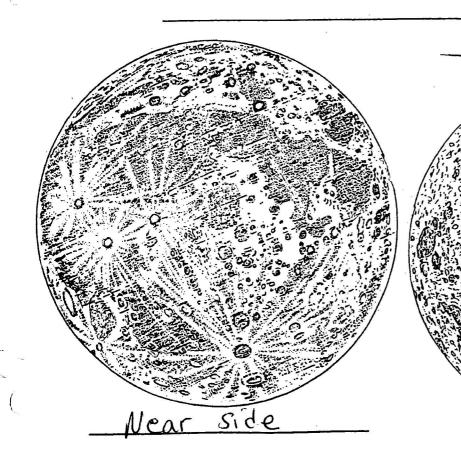


Astronomy = 30 2. Maria - appear as the "dark areas" on the moon's surface; once thought to be "_____Seas Extensive, circular, flat/smooth areas, or plainsresulted from lava flows during a much earlier period of the moon's evolution. Examples: Sea of Tranguility - Mare Tranguillitatis Sea of Showers - Mare Imbrian Sea of tears - Mare Crisium Sea of clouds - Mare Nubium $R_{\alpha\gamma}$ - appear as "bright streaks" that 3. radiate from certain craters. Consist of shattered debris that was splashed out by the impact of meteors that formed the craters. Highlands - appear as the "light areas" 4. on the moon's surface. Consist of <u>Craters</u> and mantains Carpethian, Apennine, Caucasus, Pryrenes D. THE MOON'S REVOLUTION 1. Period of Revolution a. 1 month or SCALE b. 29 1/2 days 2. The moon revolves around Earth in an _____elliptical orbit, and Earth is at one _____foci 3. This causes the moon's apparent diameter / Size to change in a <u>cyclic</u> manner.



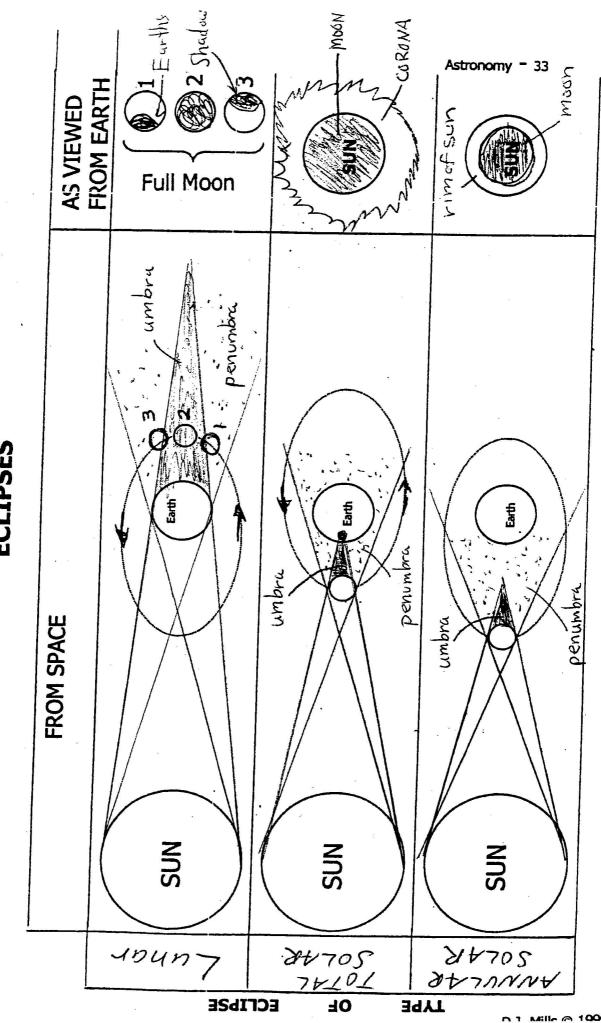


- Astronomy 32 c. (1) <u>Wanning</u> -the decreasing of the moon's visible illuminated surface; from <u>Full</u> moon to <u>New</u> moon. (2) <u>Waxing</u> -the increasing of the moon's visible illuminated surface; from <u>New</u> moon to <u>Full</u> moon.
- F. THE NEAR AND THE FAR SIDE OF THE MOON
 - 1. <u>Near Side</u> the side of the moon that always faces Earth. It is nearly half highlands and half maria.
 - 2. <u>Fac Side</u> -the side of the moon that never faces Earth. It is mostly highlands/craters.
 - 3. The same side of the moon (the near side) always faces Earth because: <u>the moon's period of rotation</u> <u>equals its period of revolution</u>



D.J. Mills © 1999

ar side



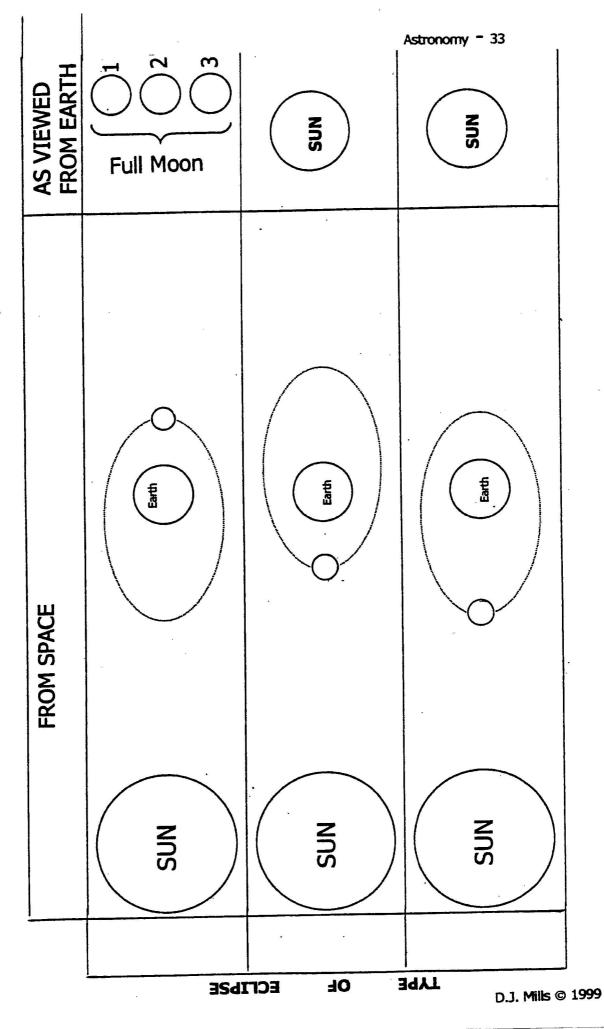
ECLIPSES

1000 A 1000

ECLIPSES

(

1-1

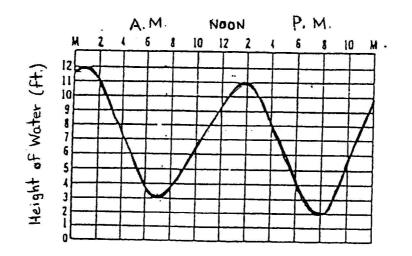


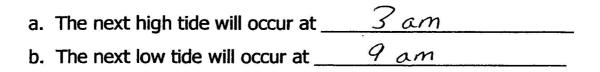
Astronomy - 34 Totality total eklipse of the sun H. TIDES 1. Tides are the periodic rising and falling of the oceans a. Caused by the moon's gravity b. Affected by Earth's _____rotation 30

2. The period from high tide to high tide is normally about

12 hours and 25 minutes. It is a cyclic change.

((





Neap tide - when the Sun, Earth and Moon form a right angle (first quarter phase and third quarter phase) causing moderate tides

Spring tide - when the Sun, Earth, and Moon form a straight line (new moon phase and full moon phase) causing high tides to be a bit higher and low tides to be a bit lower.

IX EARTH'S PLACE IN THE UNIVERSE

A. Light Year 1. the distance light travels in one tear 6,000,000,000,000 2. 3. The speed of light is 186 000 miles per second. a. The time is takes sunlight to reach each planet: Not to Scale SUN 3.2 min Zhr 40min 5hr. 30 43.3 mm min 484 million 1784 million 3675 million 36 million 93 million 10min 12.7m Zomin 6 min 142 million 67 million 887 million 2795 million b. The time it takes sunlight to reach: (1) the nearest star <u>Alpha Centauri</u> = 4.3 years (2) the brightest star Sirius = 8.6 years (3) the nearby Andromeda Galaxy _ Z, OOO, OOO years **B.** Galaxies billions of stars held together 1. by gravity 2. Shape of galaxies: a. _____Spiral D.J. Mills © 1999

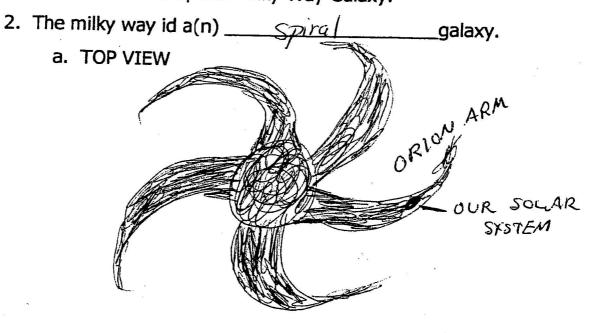
Astronomy - 37

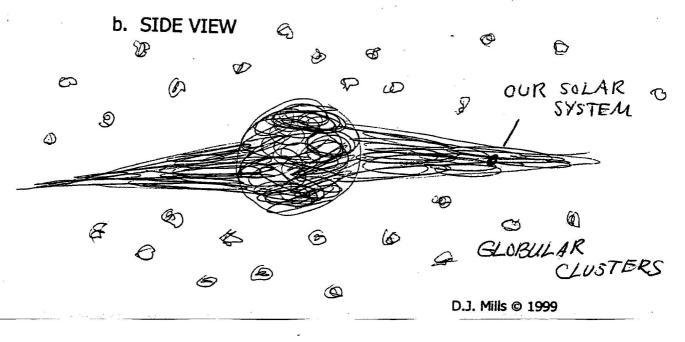
b. <u>elliptical</u> c. <u>irregular</u>

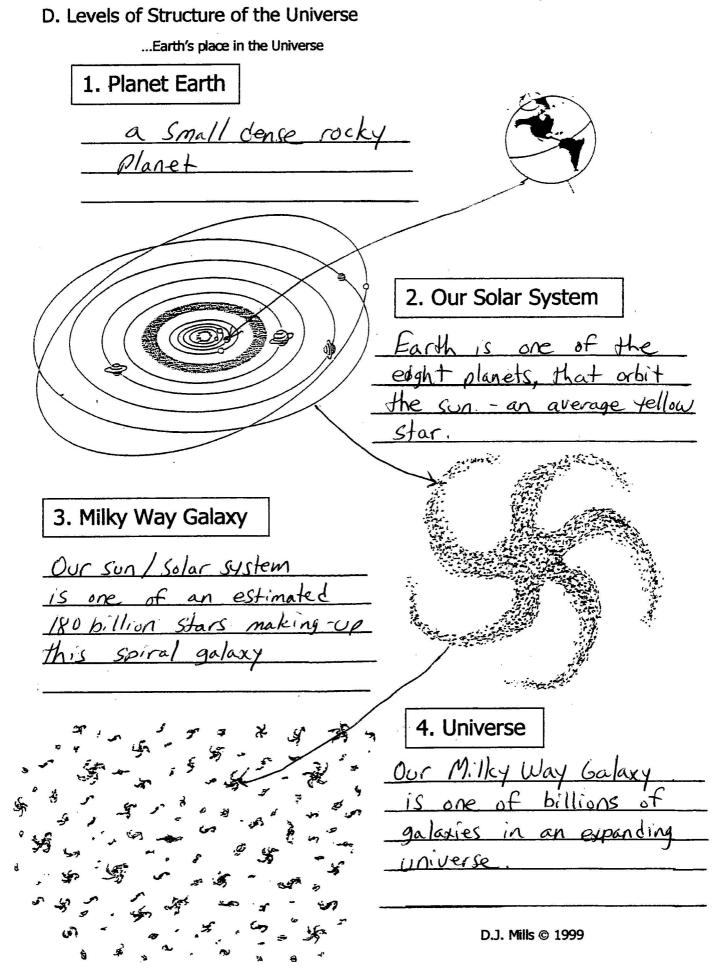




- C. The Milky Way Galaxy
 - stars that make-up the Milky Way Galaxy.







Astronomy ~ 39

E. Models of the Universe

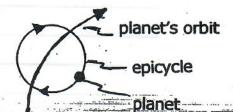
1. Geocentric Model

- a. About 2000 years ago, the Greek astronomer, Claudius Ptolemy developed a detailed model of the universe based on the idea of revolving spheres.
- b. In this model of the universe, <u>Earth</u> was the center, and all heavenly bodies moved around <u>Earth</u> in <u>Perfect circles</u>
- c. Ptolemy's <u>geocentric</u> model, as illustrated on the next page, can be summarized as follows:
 - (1.) <u>Earth</u> is located in the <u>center</u> and does not move.
 - (2.) The <u>Stars</u> are located on a transparent sphere that rotates once each day from east to west around Earth.
 - (3.) The <u>Sun</u>, the <u>mound</u> and each <u>plane</u> are carried by separate spheres of different sizes. These spheres also rotate from east to west around Earth.

(However, they rotate at slightly slower speeds than the sphere of stars and therefore have a general eastward drift relative to the stars. This explained the yearly cycle of nighttime stars.)

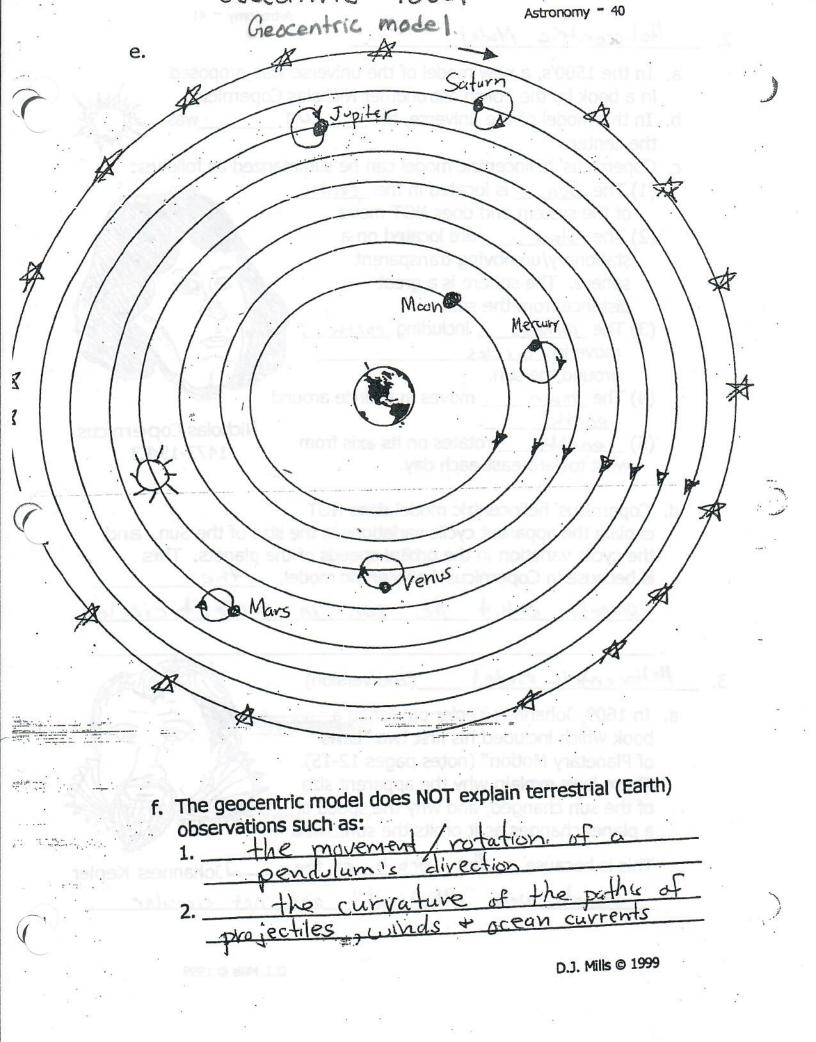
(4.) Each planet is located on an "epicycle"
(or epicircle) that also rotates. So as each planet moves around Earth on its sphere,
it is also moving or rotating on its epicycle.

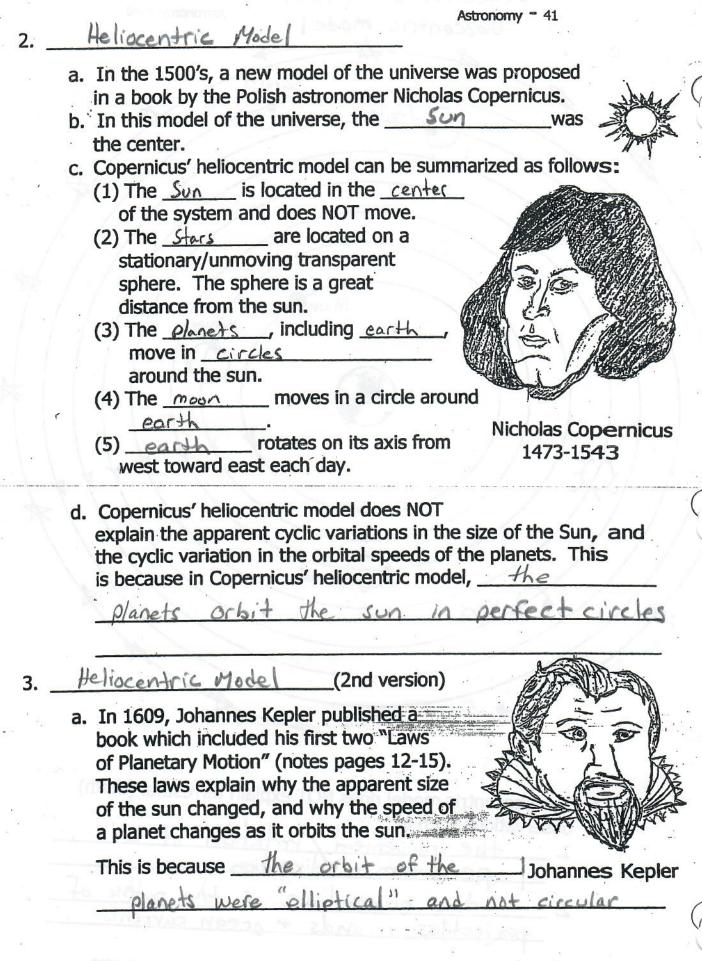
Claudius Ptolemy 100-178 A.D.



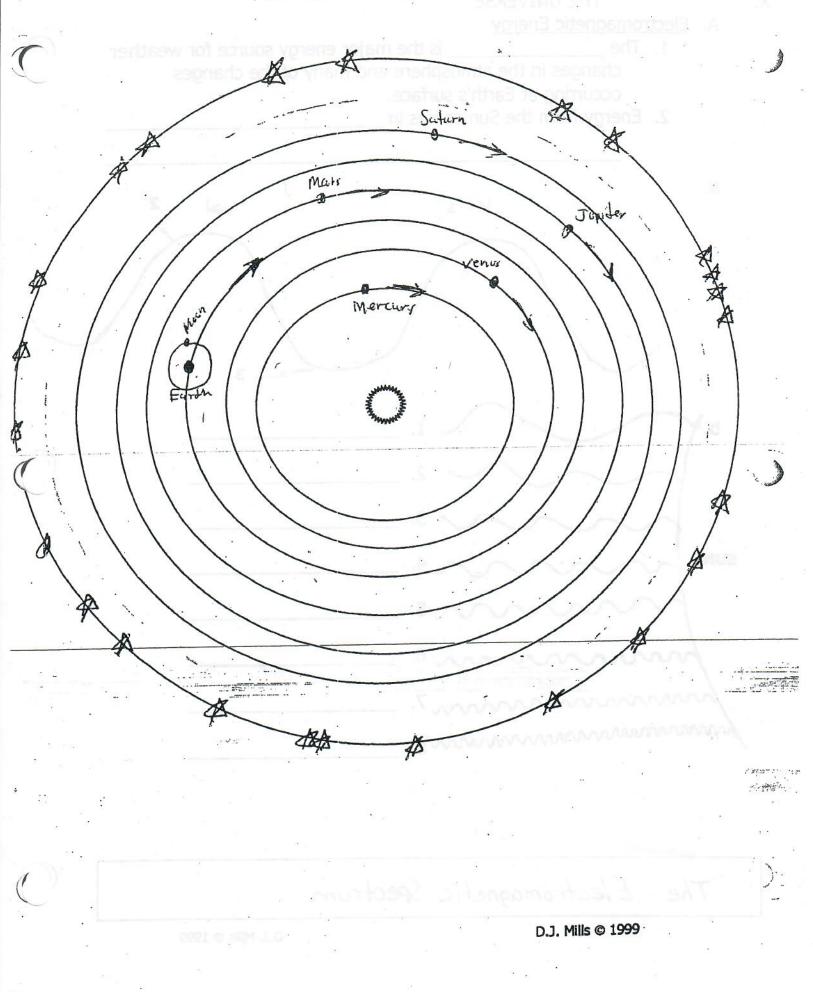
(This explained the strange "retrograde motion" of the planets relative to the background stars. That is, the planets seemed to move backward compared to the stars when you observed them for several weeks)

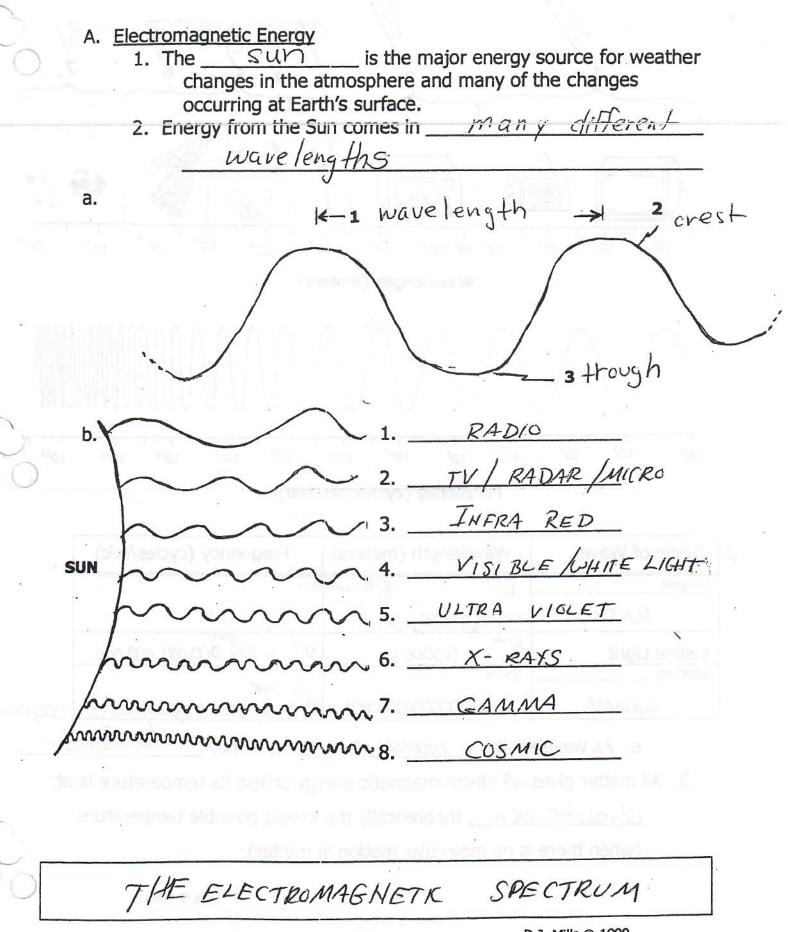
So obvious

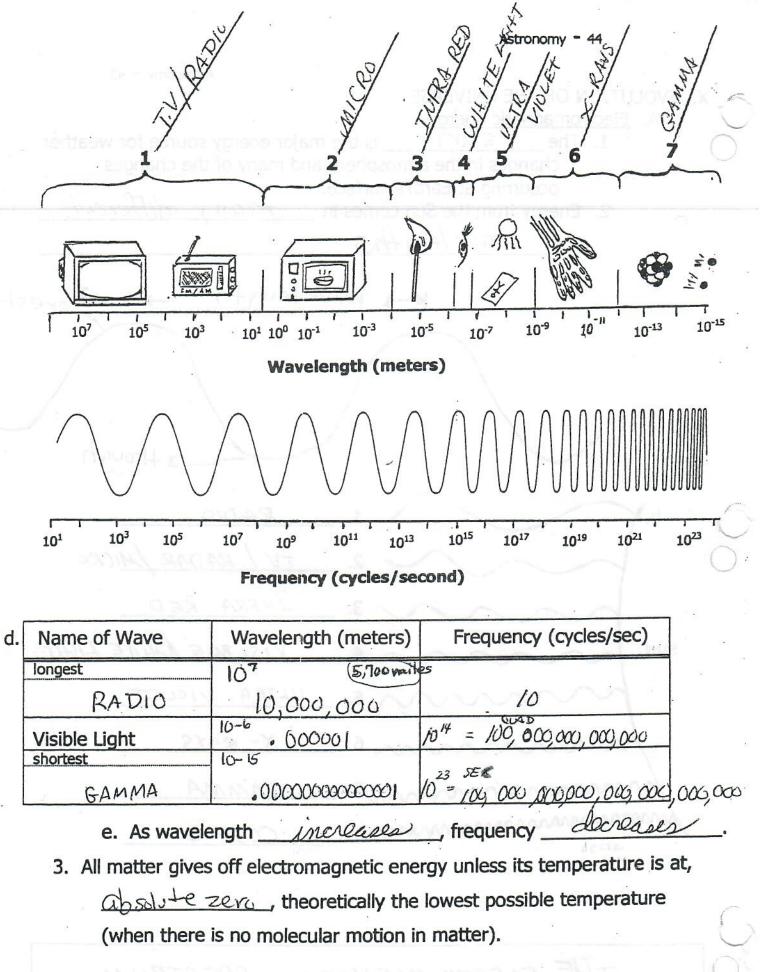




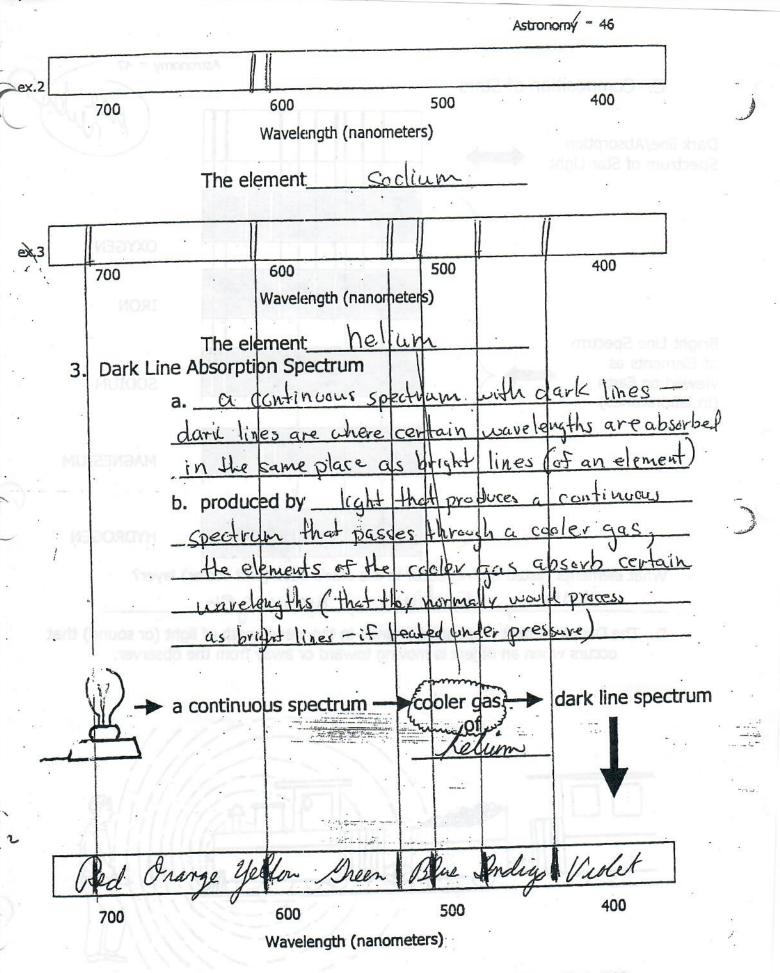
Heliocentric model Astronomy - 42



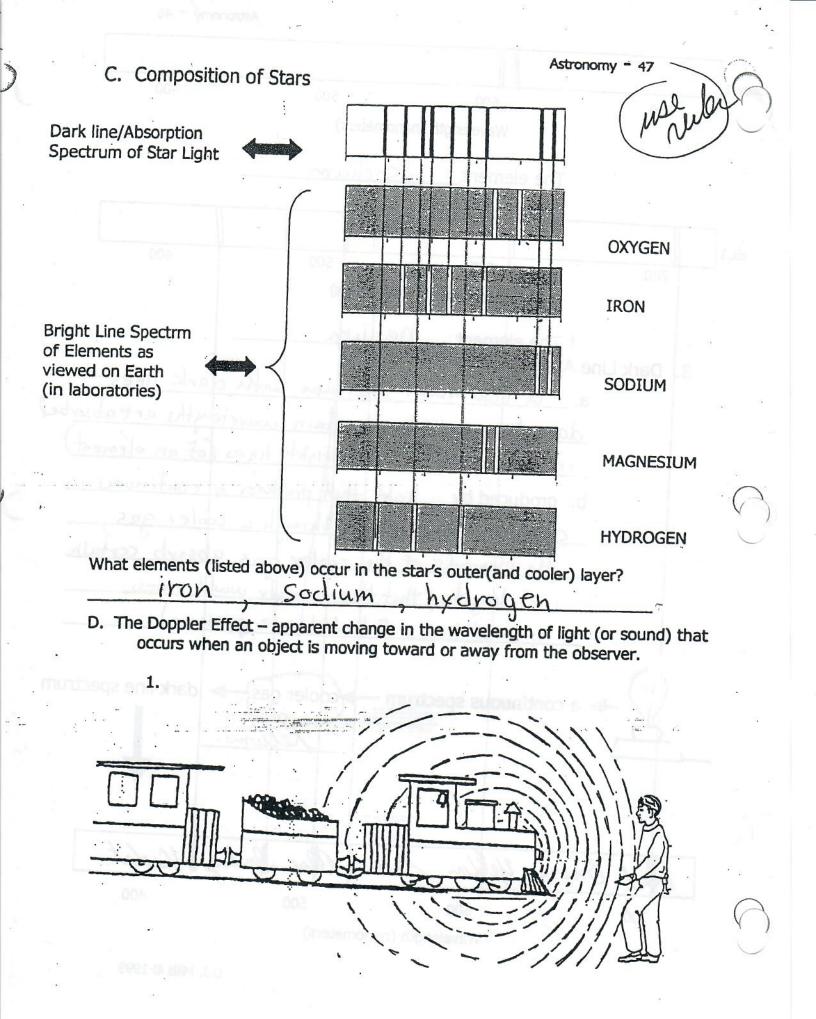




Astronomy = 45 B. Kinds of Spectra 1. White/Visible Light Red Ovange White Yellow Green Light minnin Blue Indige Violen 2. Continuous Spectrum a. unbroken band of colors b. <u>Contains all wavelengths</u> 1. glowing solid (light bulb filament) c. produced by motten lava 2. glowing liquid 3. glawing gas under pressure kompressel (star inter Mellow Green Blue Indigo Violo Travas 400 700 Wavelength (nanometers) 3. Bright Line Spectrum a. different wavelengths which appear as bright lines at different places on the spectral field. ann unique b. each element / atom has it. bright line Spectra finger c. produced by ____ chemica elemen glowing Ŷ Form 9 ex.1 400 700 500 600 Wavelength (nanometers) The element Mercur D.J. Mills © 1999

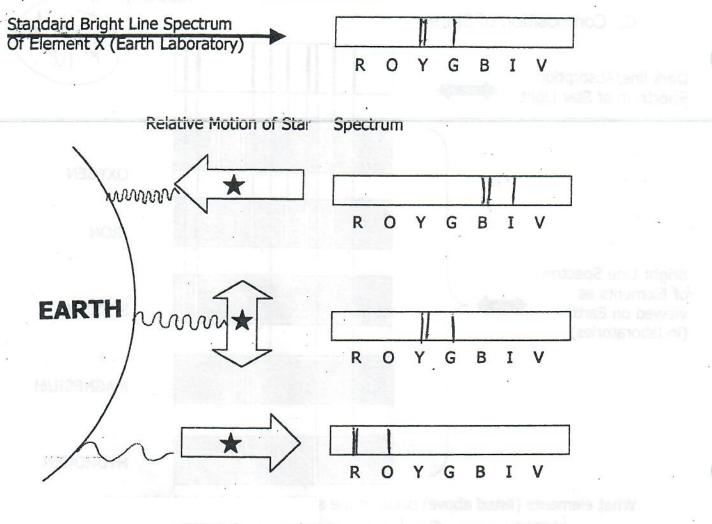


8_01



Astronomy = 48

2. Red Shift - Blue Shift Motion



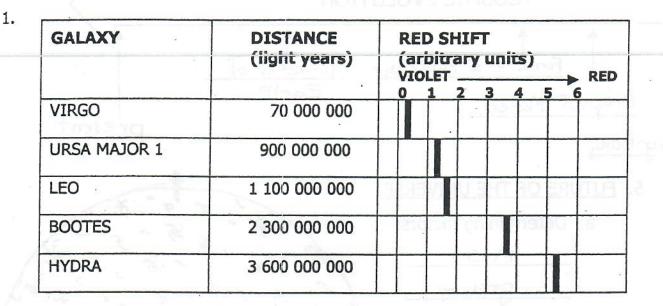
a. In the late 1920's, Edwin Hubble discovered that **ALL** galaxies were "red-shifted" when viewed from Earth.

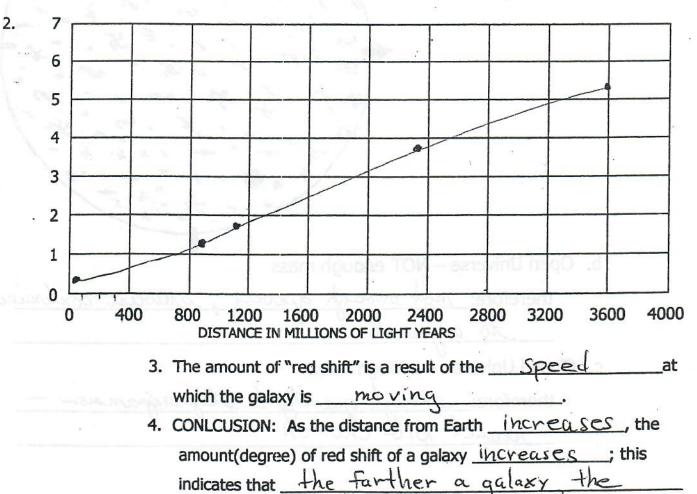
This meant that all galaxies were <u>moving away</u> <u>From Earth and each other</u>, and thus, the universe must be <u>expanding</u>

c. An echo of background radiation can still be detected by radio telescopes.

Astronomy = 49

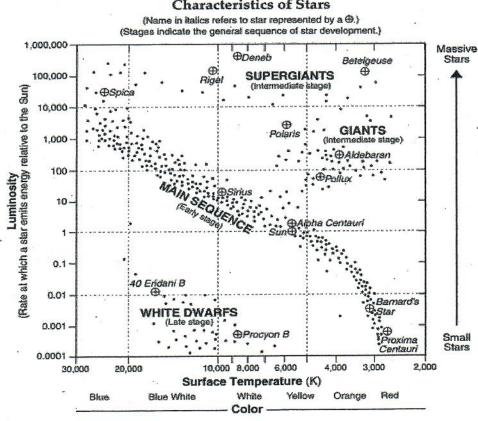
d. The data table below shows the distances of some galaxies from Earth. It also shows the "red shift" of each galaxy and how many units of red shift are exhibited by each galaxy.





faster it is moving.

and heat.



Characteristics of Stars